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Matched Subspace Detectors for Stochastic Signals*

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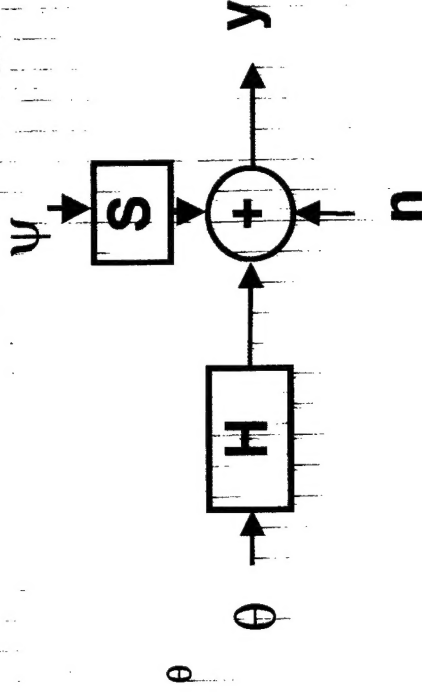
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Problem Statement

- The goal is to design detectors for stochastic signals or second-order signals.
- Extension of the first-order matched subspace detectors of Scharf and Friedlander.
- It is assumed that interference is nulled prior to processing by projecting the data into the space orthogonal to the interference subspace.
- We assume various states of knowledge about the parameters σ^2 and β .



$$n \sim \text{CN}(0, \sigma^2 I)$$

$$\theta : f(\theta; \beta) \text{ ex. } \theta \sim \text{CN}(0, R_{\theta\theta})$$



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$$\begin{aligned} Z &= (H^*(I-P_S)H)^{-1/2} H^*(I-P_S) Y \\ &= (G^*G)^{1/2} G^* Y \end{aligned}$$

$$W = A^* y$$

Hypotheses

- The "noise" vector w is distributed as a white complex Gaussian vector regardless of which hypothesis is in effect.

$$f(w) = \frac{1}{(\pi\sigma^2)^N} e^{-\frac{w^*w}{\sigma^2}}$$

- Define $\phi = (G^*G)^{1/2}\theta$.

$$f(z | \phi) = \frac{1}{(\pi\sigma^2)^p} e^{-\frac{1}{\sigma^2} \|z - \phi\|^2}$$

- When signal is present the data vector z is distributed:

$$f(z | \phi \neq 0) = \frac{1}{(\pi\sigma^2)^p} e^{-\frac{z^*z}{\sigma^2}}$$

- When signal is not present the data vector z is distributed:



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Likelihood Ratio



- For now, assume that the noise power σ^2 is known.
- In this case the vector w is common to both hypotheses and is of no use.
- The *conditional* likelihood ratio is then

$$\begin{aligned} l(z \mid \phi; \sigma^2) &= \frac{f(z \mid \phi; \sigma^2)}{f(z \mid \phi=0; \sigma^2)} \\ &= \exp\left(\frac{z^* z}{\sigma^2}\right) \times \exp\left(-\frac{1}{\sigma^2} \|z - \phi\|^2\right) \end{aligned}$$



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Unconditional Likelihood Ratio



- The unconditional likelihood ratio can be written as

$$l(z; \sigma^2, \beta) = \exp\left(\frac{z^*z}{\sigma^2}\right) \times \int \exp\left(-\frac{\|z-\phi\|^2}{\sigma^2}\right) f_\phi(\phi; \beta) d\phi$$

- The log-likelihood ratio becomes

$$s(z; \sigma^2, \beta) = \frac{z^*z}{\sigma^2} + \ln \int \exp\left(-\frac{\|z-\phi\|^2}{\sigma^2}\right) f_\phi(\phi; \beta) d\phi$$

$$= \frac{y^*P_{GY}}{\sigma^2} - p_r(z; \sigma, \beta)$$

Matched Subspace
Detector

Resolution penalty



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Resolution Penalty



- The resolution penalty occurs because we presume to know something about the coordinate vector θ .
- If z is far from the "favored" orientation defined by θ then the penalty is larger than if the converse were true.

$$p_r(z; \sigma^2; \beta) = -\ln \int \exp\left(-\frac{\|z - \phi\|^2}{\sigma^2}\right) f_\phi(\phi; \beta) d\phi$$



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Gaussian Coordinate Vectors



- Suppose $\phi \sim \text{CN}(0, R_{\phi\phi})$.
- Write the eigenvalue decomposition of $R_{\phi\phi}$ as:

$$R_{\phi\phi} = (G^*G)^{1/2} R_{\theta\theta} (G^*G)^{1/2} = V D^2 V^*$$

$$V = [v_1 \ v_2 \ \dots \ v_p]; \text{ unitary}$$

$$D^2 = \text{diag}[\beta_1^2, \beta_2^2, \dots, \beta_p^2]$$
- Define the resolved signal-plus-noise to noise ratios:

$$r_i = 1 + \frac{\beta_i^2}{\sigma^2}$$



Gaussian Penalty Term



- After some algebra the penalty term can now be written as

$$p_r(\mathbf{z}; \sigma^2, \beta^2) = -\ln \int \exp\left(-\frac{\|\mathbf{z} - \phi\|^2}{\sigma^2}\right) \frac{1}{\pi^p \det(R_{\phi\phi})} \exp(-\phi^* R_{\phi\phi}^{-1} \phi) d\phi$$

$$= \sum_{i=1}^p \ln(r_i) + \sum_{i=1}^p \frac{(\mathbf{z}^* P_{V_i} \mathbf{z} / \sigma^2)}{r_i}$$

- This result implies that if the estimated signal-plus-noise to noise ratio $(\mathbf{z}^* P_{V(i)} \mathbf{z} / \sigma^2)$ in the resolved subspace defined by \mathbf{v}_i greatly exceeds r_i , then the penalty is large because of this mismatch.



Unknown Signal Power and Orientation



- Suppose that when signal is present we do not know $R_{\phi\phi}$.
- Recall the penalty term is

$$p_T(\mathbf{z}; \sigma^2, \beta^2) = -\ln \int \exp\left(-\frac{\|\mathbf{z}-\phi\|^2}{\sigma^2}\right) \frac{1}{\pi^p \det(R_{\phi\phi})} \exp(-\phi^* R_{\phi\phi}^{-1} \phi) d\phi$$

$$= \sum_{i=1}^p \ln(r_i) + \sum_{i=1}^p \frac{(\mathbf{z}^* P_{V_i} \mathbf{z} / \sigma^2)}{r_i}$$

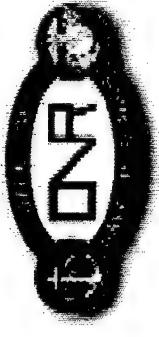
- The estimates of the signal-plus-noise to noise ratios are

$$r_i = \max(1, \mathbf{z}^* P_{V_i} \mathbf{z} / \sigma^2)$$

- We assume that $r_i \geq 1$ in the sequel.



Estimating Orientation



- The estimates of r_i in the previous slide depend on the orientation of the vectors v_i .

- We want to minimize

$$\prod_{i=1}^p \frac{z^* P_{v_i} z}{\sigma^2}$$

- We must also satisfy the constraints

$$\sum_{i=1}^p \frac{z^* P_{v_i} z}{\sigma^2} = \frac{z^* z}{\sigma^2}$$

$$r_i = \frac{z^* P_{v_i} z}{\sigma^2} \geq 1$$



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Intermediate Orientation Solution



- The solution to this optimization problem is

$$r_i = \frac{z^* P_{V_i} z}{\sigma^2} = 1 \quad \text{for } i = 1, 2, \dots, p-1$$

$$r_p = \frac{z^* P_{V_p} z}{\sigma^2} = \frac{z^* z}{\sigma^2} - (p-1)$$

- The question remains: Is there a decomposition of $\langle G \rangle$ that has the above properties?
- The answer is yes.



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Orientation Solution



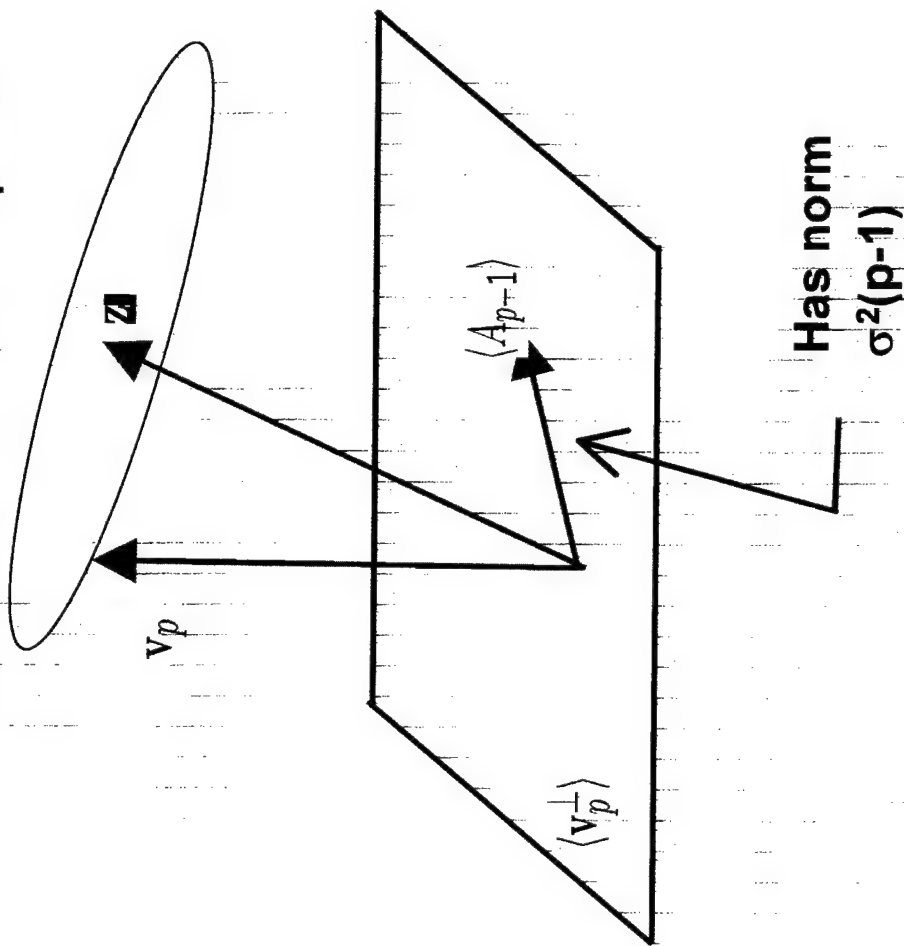
- Solve for v_p first.
- Choose a v_p on the spherical invariance set defined by

$$\frac{z^* P_{v_p} z}{\sigma^2} = \frac{z^* z}{\sigma^2} - (p-1).$$

- Repeat this procedure in the spaces

$$\langle A_{p-1} \rangle, \langle A_{p-2} \rangle, \dots, \langle A_1 \rangle$$

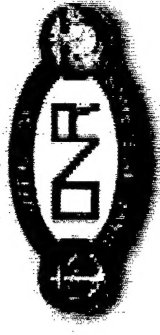
Great circle on invariance sphere



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Compressed Likelihood



- Compressing the likelihood ratio with this solution gives the statistic

$$s(z; \sigma^2, \hat{R}_{\phi\phi}) = \frac{y^* P_{HY}}{\sigma^2} - \left[\ln \left(\frac{y^* P_{HY}}{\sigma^2} \right) - constants \right]$$

- This statistic is a monotonic function of the matched subspace detector. We can therefore use the MSD as the detection statistic

$$s = \frac{y^* P_{HY}}{\sigma^2}$$

- Then the result for 2nd-order models is the same as for 1st-order models.



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Unknown Noise Power



- In the case of unknown noise power the GLRT detector can be written as a sum of the CFAR matched subspace detector and a penalty term

$$s(\mathbf{z}; \hat{\sigma}^2, \hat{R}_{\phi\phi}) = \ln(1 + \tilde{s}) + [\ln(\tilde{s}) - \text{constants}]$$

- We can equivalently use the statistic

$$\tilde{s} = \frac{\mathbf{y}^* P_{HY}}{\hat{\sigma}^2}; \quad \hat{\sigma}^2 = \frac{1}{N-p} \mathbf{y}^* (\mathbf{I} - P_H) \mathbf{y}$$

- These detectors are identical to the 1st-order results.



Rank-One Assumptions

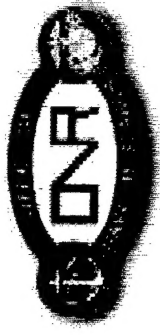


- Here we assume that the signal subspace is rank-one.
- The complex-valued signal amplitude is written in polar form
$$\theta = M e^{j\phi}$$
- Assume that the phase and magnitude are uncorrelated and that the phase is uniformly distributed over $[0, 2\pi)$.
- Assume that the signal magnitude has a generalized Rayleigh distribution

$$f_M(M) = \frac{2M}{\beta^2} \left(\frac{M^2}{\beta^2} \right)^{L-1} \frac{e^{-M^2/\beta^2}}{L!}$$



Detectors with Known Noise



- $L=0$. This is the previous results with complex Gaussian amplitudes.

- $L \neq 0$. The penalty function is

$$p_r = (L+1) \ln(r) + \frac{(y^* P_h y / \sigma^2)}{r} - \ln \left[\sum_{k=0}^L \frac{\text{binom}(L,k)}{k!} \left(\frac{(y^* P_h y / \sigma^2)(r-1)}{r} \right)^k \right]$$

- Minimize the penalty term with respect to $r=1+\beta^2/\sigma^2$.
- Compress the likelihood function with this term to obtain

$$s = \frac{y^* P_h y}{\sigma^2} - p_r(\hat{r}).$$



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